

National Aeronautics and Space Administration



# SLS-RFM\_23-04 Performance Validation of Bi-Phase Filtering for Category A Space-to-Earth Links

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# Background

- During Fall 2022 meeting, ESA presented input paper SLS-RFM\_22-10 on bi-phase filtering for Category A Space-to-Earth links so that the spectrum is compliant to the SFCG mask for symbol rates  $< 300\text{ks/s}$ .
- The recommendation is to apply a 3<sup>rd</sup> order Butterworth filter before phase modulation of the baseband signal (referred as Option 2).
- The goal of this effort is to verify ESA's spectral and BER results by MATLAB simulation which serve as a means of independent validation required by CCSDS.

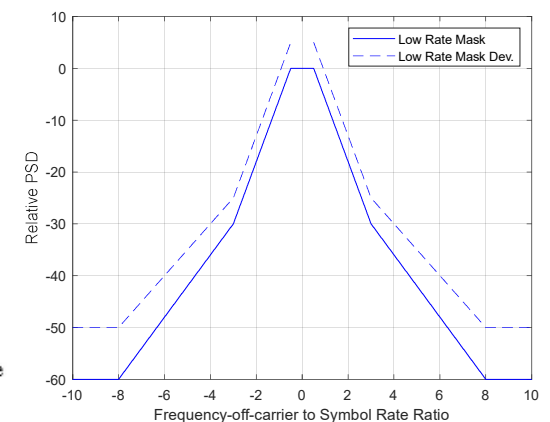
# SFCG 21-2R4 Guideline for Bi-Phase

## RECOMMENDS

1. that space agencies use the most bandwidth efficient modulation schemes practicable for their missions;
2. that, PCM/PM/Bi-phase or PCM/PM/NRZ modulation only be used when a carrier component is technically necessary and for symbol rates below 2 Ms/s.
3. that the emitted spectrum<sup>3,4</sup> for all Space Science Services projects that will utilize space-to-Earth link frequency assignments in the bands 2200–2290 MHz, 8025–8400 MHz and 8450–8500 MHz, adhere to the low rate spectral emission mask of Figure 1 for symbol rates below 2 Ms/s and to the high rate spectral emission mask of Figure 1 for symbol rates equal or above 2 Ms/s;
4. that the emitted spectrum<sup>3</sup> for all Space Science Services projects designed for launch after 2020 that will utilize space-to-Earth link frequency assignments in the 25.5–27.0 GHz band and for channel symbol rates<sup>5</sup> equal or above 10 Ms/s, adhere to the high rate spectral emission mask of Figure 1;
5. that transmissions that include a ranging signal be exempt from the spectrum masks in Fig 1;
6. that PCM/PSK/PM transmissions in accordance with REC SFCG 21-3 be exempt from the spectrum masks in Fig 1.

<sup>3</sup> Measured relative to the peak of the telemetry spectrum and excluding the residual carrier as well as all spurious emissions.  
<sup>4</sup> PCM/PM/Bi-phase emissions with symbol rates up to 300 ks/s may deviate from the low rate mask by up to 5 dB in the slope region and up to 10 dB in the plateau region, and in the transition between the two regions.

Interpretation:  
1. Relaxation on the low rate mask is allowed for bi-phase with symbol rates up to 300ks/s.



# Mathematical Model for Unfiltered and Filtered Bi-Phase Signal with Residual Carrier

- Bi-phase with no filtering:

$$s(t) = e^{jm \underbrace{\sum_k a_k p(t-k/R_s)}_{\text{This term has power} = 1}}$$

where

$m$  – input modulation index

$a_k$  –  $k^{\text{th}}$  data value (-1 or 1)

$p(t)$  – bi-phase pulse shape

$R_s$  – symbol rate (before bi-phase converter)

- Bi-phase with filtering before PM:

$$s(t) = e^{jm \underbrace{\sum_k a_k (g * p)(t-k/R_s)}_{\text{This term has avg. power} < 1 \text{ due to filtering}}}$$

or  $\underbrace{\text{Effective mod index } m'}$

$$s(t) = e^{jm \sqrt{C} \frac{1}{\sqrt{C}} \underbrace{\sum_k a_k (g * p)(t-k/R_s)}_{\text{This term has avg. power} = 1}}$$

where

$C$  – avg power of  $\sum_k a_k (g * p)(t - k/R_s)$

$g(t)$  – spectral shaping filter

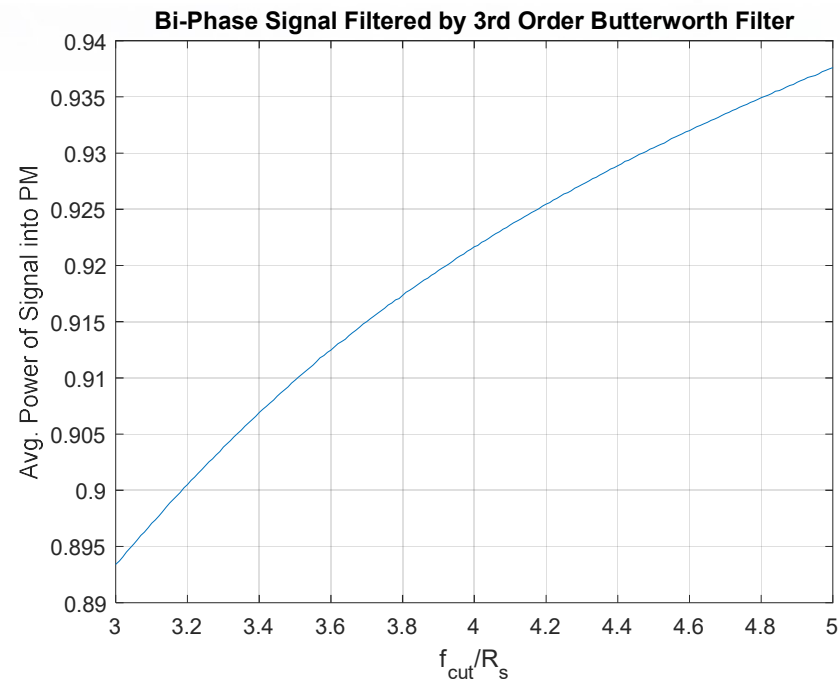
\* denotes the convolution operator

ESA has chosen 3<sup>rd</sup> order Butterworth filter for  $g(t)$  with a range of cutoff frequencies to satisfy the SFCG spectral mask for different modulation indices.

# Effective Modulation Index

- Since average power in the data modulation is reduced by the spectral shaping filter, the constant term  $C$  needs to be estimated by simulation so that input mod index,  $m$ , can be adjusted accordingly to achieve a desired “effective” mod index,  $m'$ .

$$m = \frac{m'}{\sqrt{C}}$$

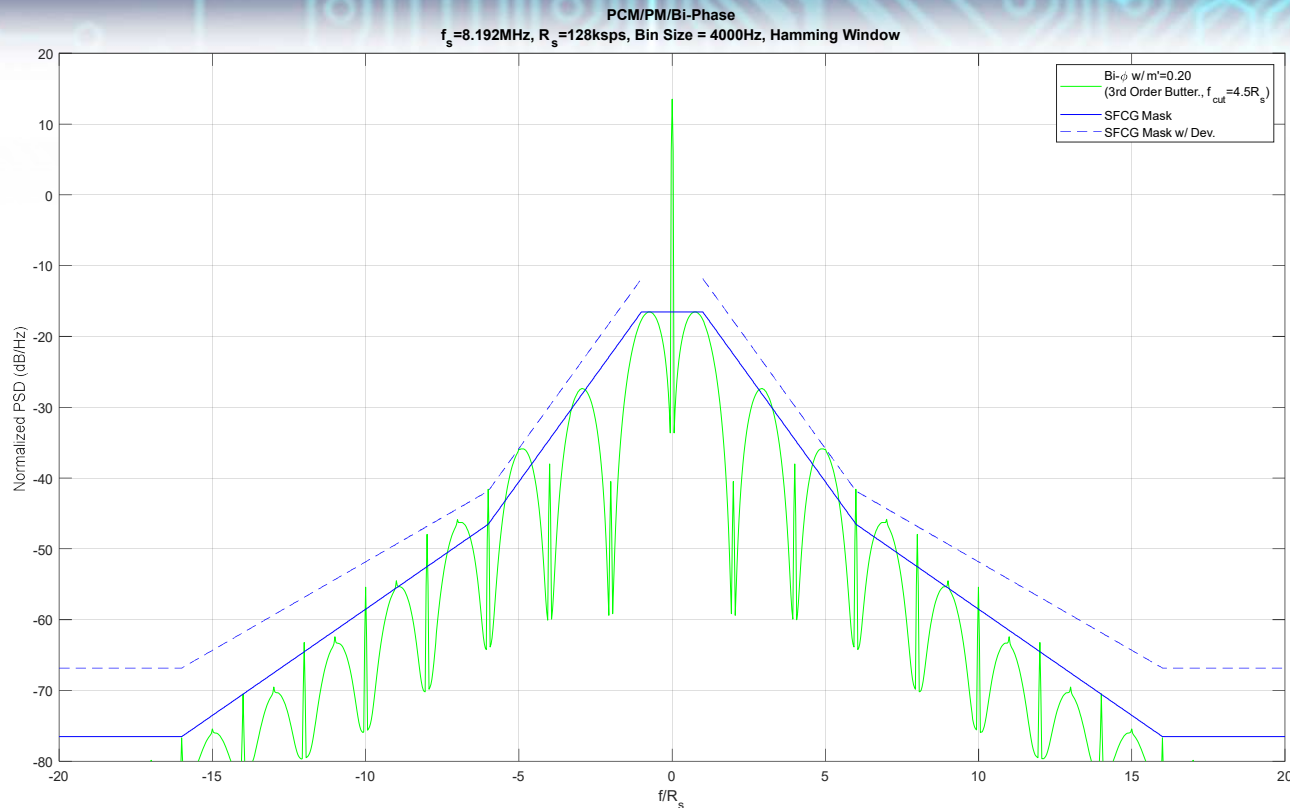




# Spectral Analysis

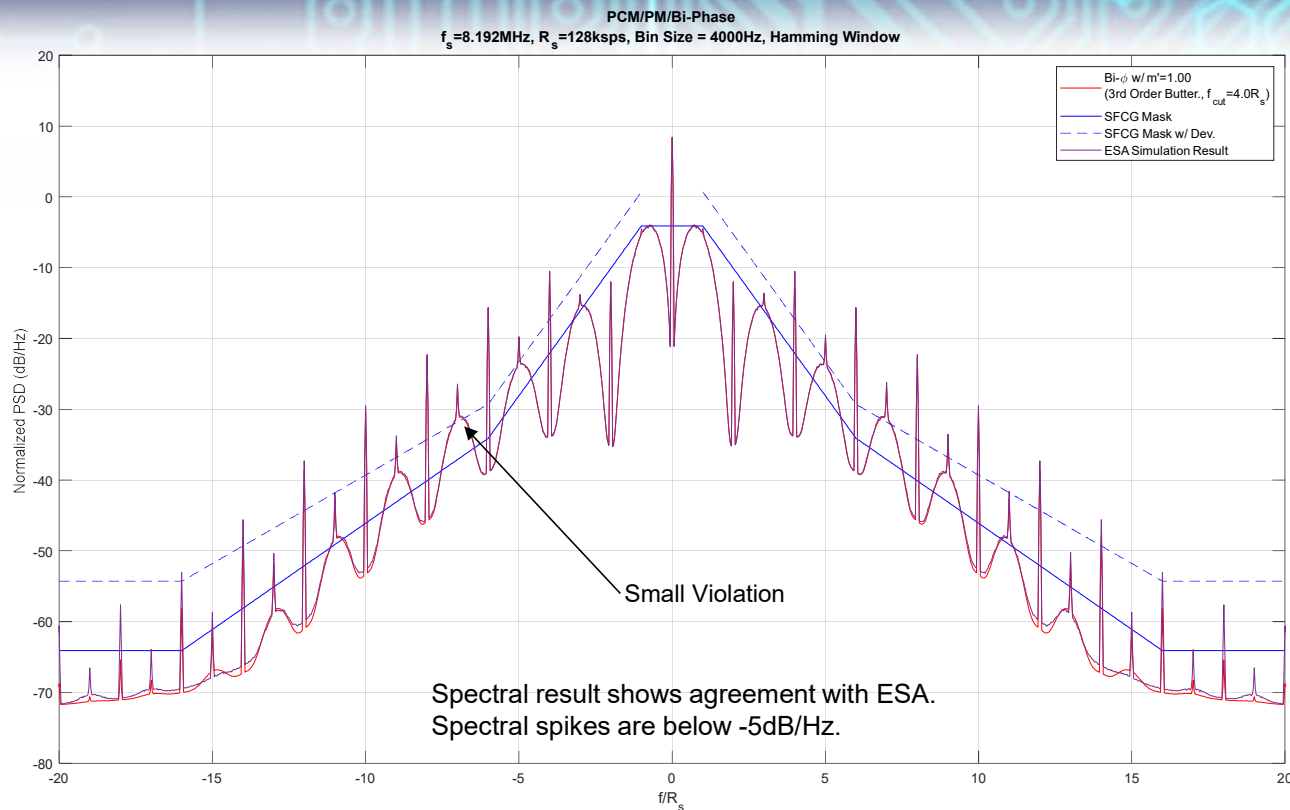
- ESA showed that  $f_{\text{cut}}$  settings of  $4.5R_s$ ,  $4.0R_s$ ,  $3.5R_s$  for  $m' = 0.2, 1, 1.25$  are sufficient to meet SFCG mask.

# Effective Mod Index 0.2 rad-pk with $f_{\text{cut}} = 4.5R_s$



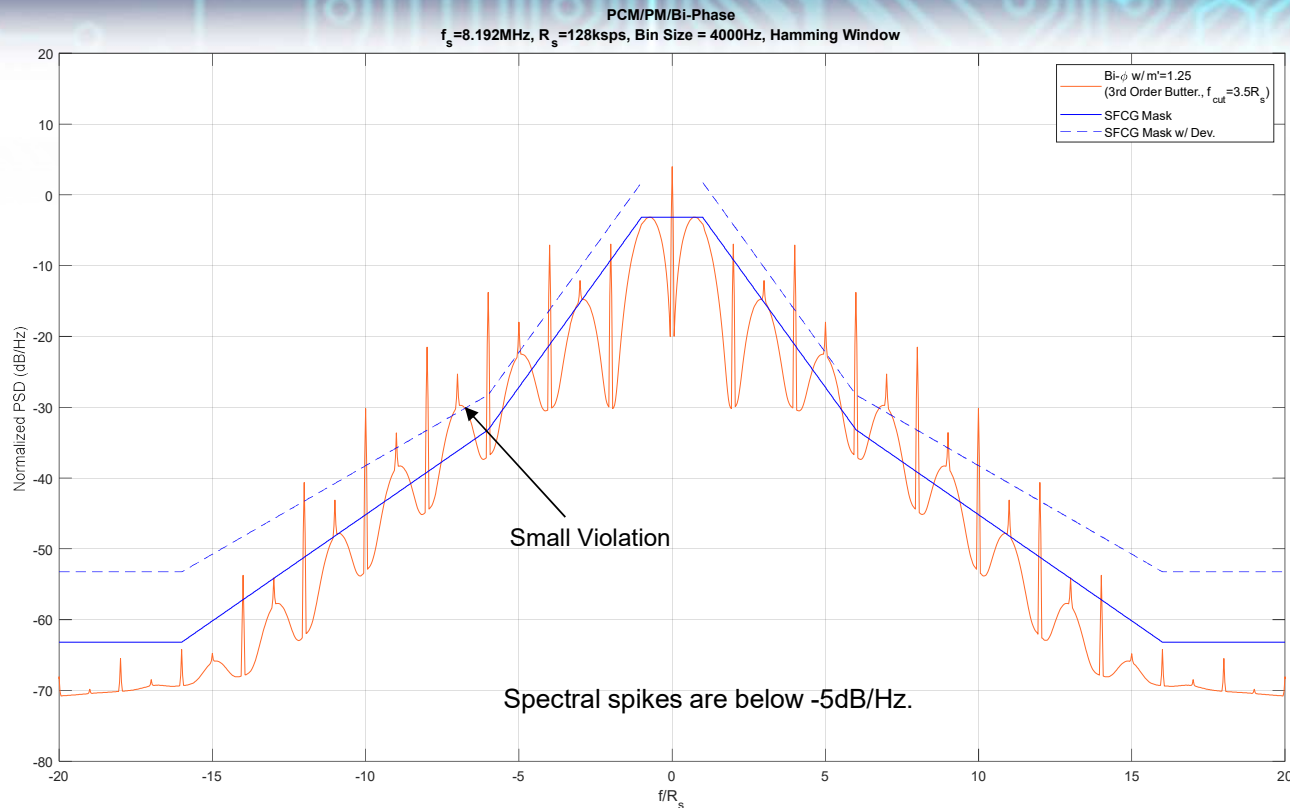
Spectrum meets SFCG mask.  
Spectral spikes are below -5dB/Hz  
(ESA's requirement for spectral spike is  
20dBc below unmodulated carrier)

# Effective Mod Index = 1.0 rad-pk with $f_{\text{cut}} = 4.0R_s$



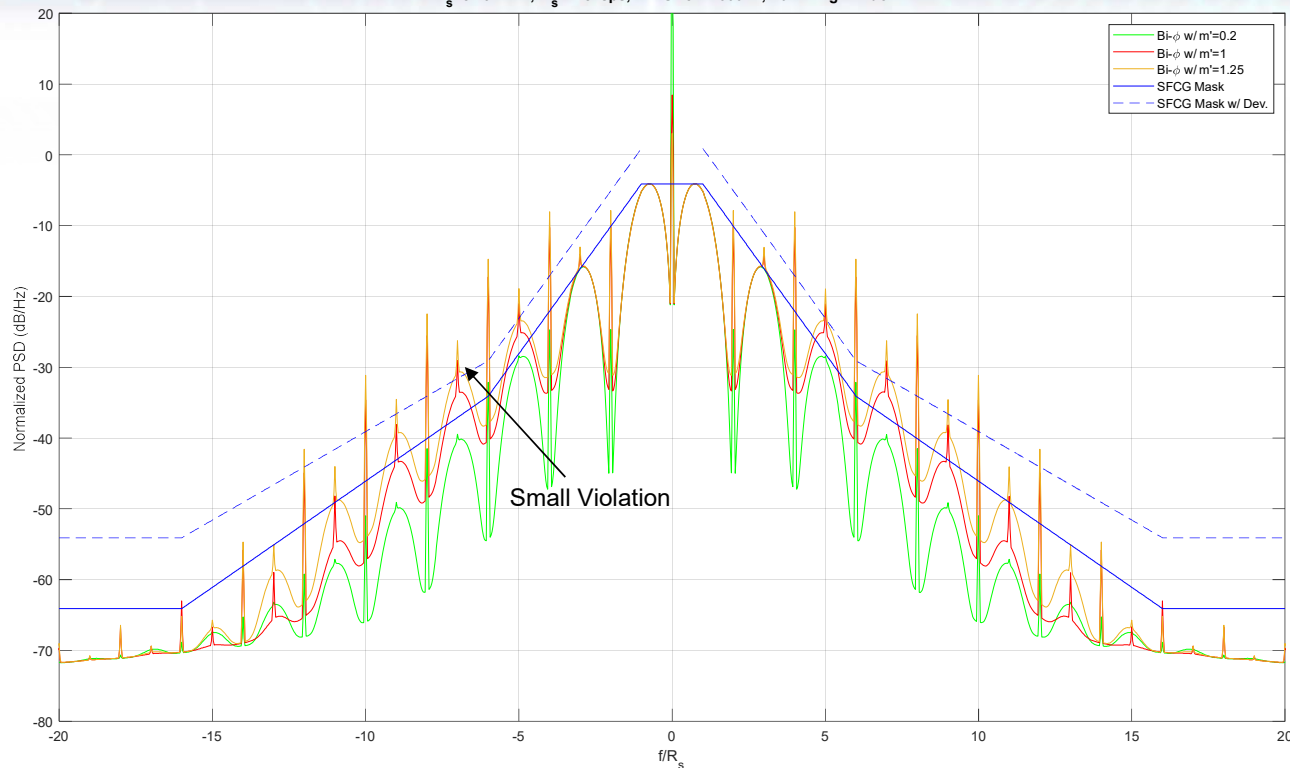


# Effective Mod Index = 1.25 rad-pk with $f_{\text{cut}} = 3.5R_s$



# Effective Mod Index = 0.2, 1, 1.25 rad-pk with $f_{\text{cut}}=3.5R_s$

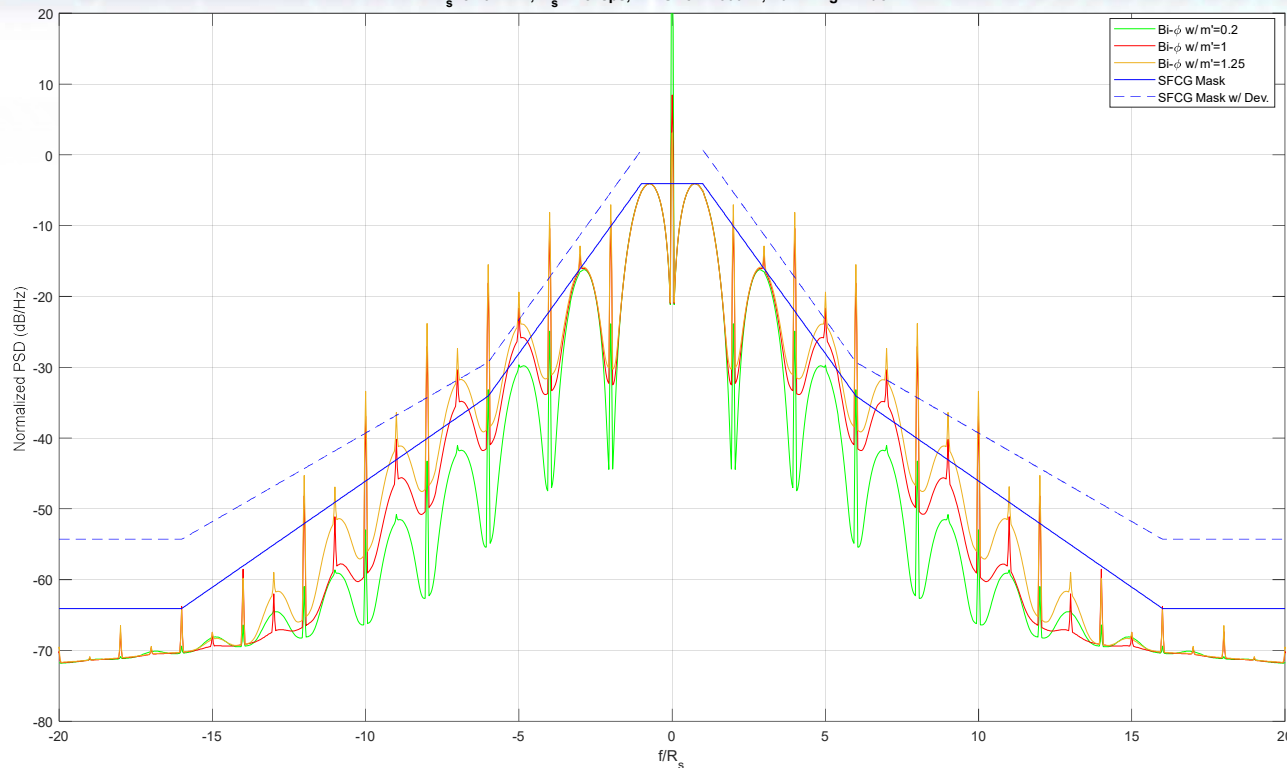
Bi-Phase w/ 3rd Order Butterworth Filter ( $f_{\text{cut}}=3.5R_s$ )  
 $f_s=8.192\text{MHz}$ ,  $R_s=128\text{kpsps}$ , Bin Size = 4000Hz, Hamming Window



Note: Plots are shifted to the same level for comparison.

# Effective Mod Index = 0.2, 1, 1.25 rad-pk with $f_{\text{cut}} = 3.3R_s$

Bi-Phase w/ 3rd Order Butterworth Filter ( $f_{\text{cut}} = 3.3R_s$ )  
 $f_s = 8.192\text{MHz}$ ,  $R_s = 128\text{ksp}$ s, Bin Size = 4000Hz, Hamming Window



Spectral results show that there is some small potential spectral violations for  $m'$  of 1.25 using  $f_{\text{cut}}$  of  $3.5R_s$  therefore we reduced  $f_{\text{cut}}$  setting to  $3.3R_s$ .

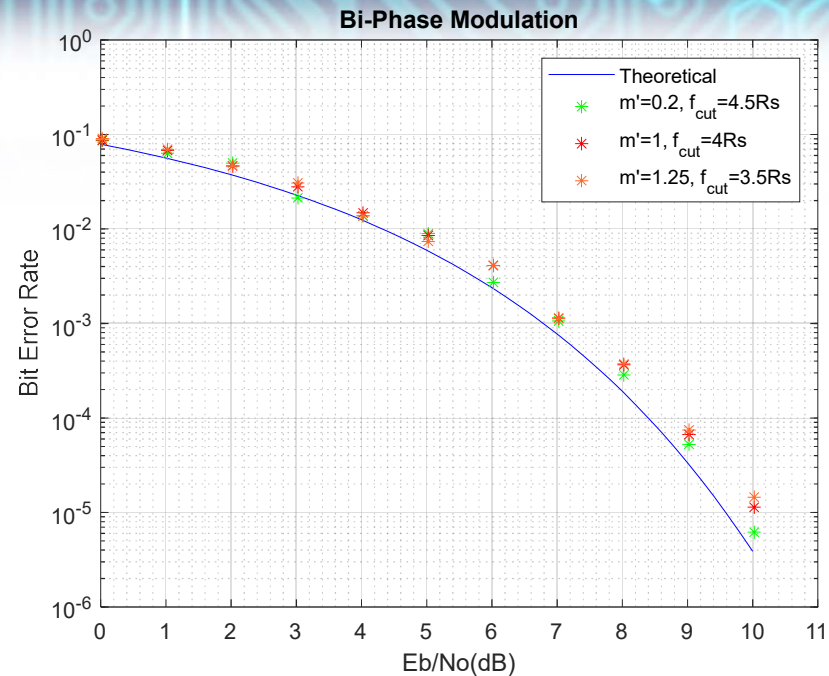
Note: Plots are shifted to the same level for comparison.



# Bit Error Rate Analysis

- Assumptions:
  - Receiver is unaware of the additional filtering and therefore uses the bi-phase pulse as the matched filter.
  - Therefore BER simulation plots are generated based on the expected  $E_b/N_0$  for unfiltered bi-phase signal with power of  $\sin^2(m')$ .
  - Channel model is AWGN.

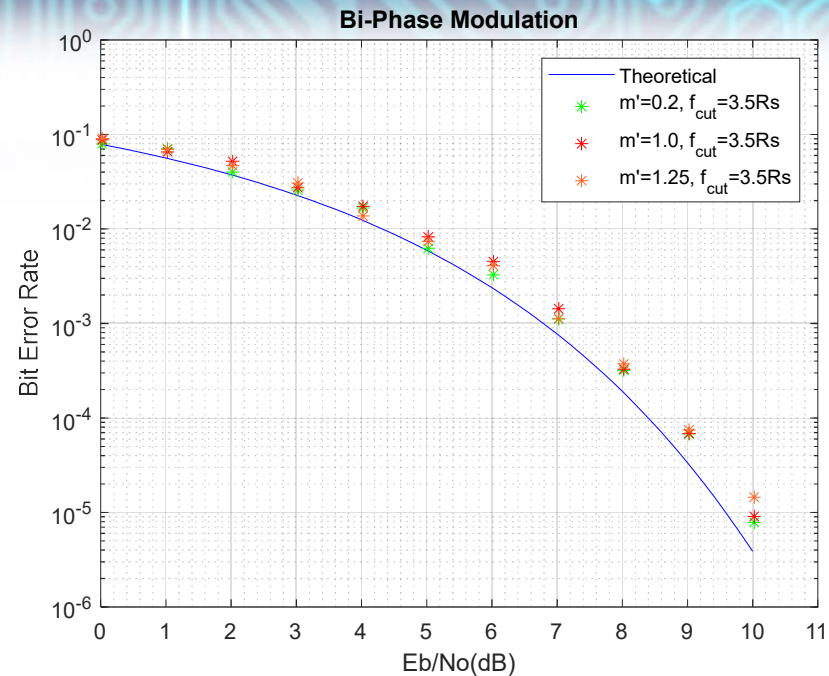
# Effective Mod Index 0.2, 1.0, 1.25 rad-pk at $f_{\text{cut}}=4.5Rs, 4Rs, 3.5Rs$



Loss @ $1e-5$	ESA's Result	NASA's Result
$m'=0.2, f_{\text{cut}}=4.5Rs$	~ 0.2 dB	~ 0.2 dB
$m'=1, f_{\text{cut}}=4Rs$	~ 0.4 dB	~ 0.5 dB
$m'=1.25, f_{\text{cut}}=3.5Rs$	~ 0.5dB	~ 0.6 dB

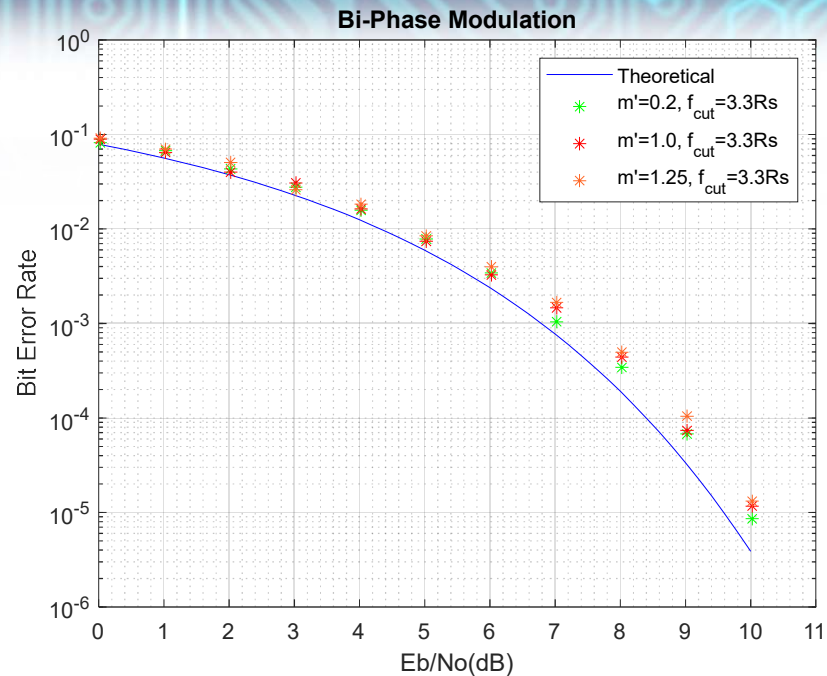
NASA's results are similar to ESA's.

# Effective Mod Index 0.2, 1.0, 1.25 rad-pk at $f_{\text{cut}}=3.5R_s$



If smallest filter cutoff frequency,  $3.5R_s$ , is used for all mod indices, implementation loss ranges from about 0.3 to 0.6dB at  $1e-5$  BER.

# Effective Mod Index 0.2, 1.0, 1.25 rad-pk at $f_{\text{cut}} = 3.3R_s$



If smallest filter cutoff frequency,  $3.3R_s$ , is used for all mod indices, implementation loss ranges from about 0.3 to 0.6dB at  $1e-5$  BER.

# Some Observations

- NASA spectral requirement is to meet both SFCG and NTIA masks.
  - We are currently trying to understand spectral regulations on the spikes.
- Section 5.6.2 of Redbook indicated that single frequency unwanted emissions be at least 60dBc below mean transmit power. Spikes shown in spectral plots do not meet this requirement.
- Per Dennis Lee's initial suggestion, it might be possible to argue that the spikes are a necessary byproduct of PM filtering, and be considered as part of the inherent modulation spectrum.
- NTIA Annex J regulations, Section J.3.4 January 2021 Edition, Table 5 is shown on right with regards to PCM/PM/Bi-Phase

K-factor specifications for filtered PCM/PM/Bi-Phase from NTIA Annex-J document:

**Table 5: K-factors for PCM/PM Bi-Phi with Filtering and without Subcarrier**

Modulation Index (m)	Filter Type	Filter Characteristics	K-factor
1.2	Square Root Raised Cosine	2000 tap; roll off factor=1	2x2.0
1.2	Butterworth	3 pole; Bandwidth Symbol Time Product (BTs)=2	2x3.0
1.2	Bessel	3 pole; Bandwidth Symbol Time Product (BTs)=2	2x3.0
$\pi/2$	Square Root Raised Cosine	2000 tap; roll off factor=1	2x2.6
$\pi/2$	Butterworth	3 pole; Bandwidth Symbol Time Product (BTs)=2	2x4.0
$\pi/2$	Bessel	3 pole; Bandwidth Symbol Time Product (BTs)=2	2x4.0

Definition for necessary BW, Bn, is not specified for bi-phase modulated signal. The closest one is for BPSK signal where  $B_n = 2 * R_s * K$ .

- Table 3.2.2.2-1 in ICSIS document specifies PCM/PM/Bi-Phase-L be used for the X-band return link when ranging is required, for symbol rates (before bi-phase converter) between 64ksps and 1.024Mpsps.
  - SFCG allows exemption for transmissions with ranging signal, but NTIA doesn't. Necessary BW needs to be increased to account for this.

**X-band Return Link (8450-8500 MHz)<sup>6</sup> (CSP-Lunar Exploration System to Earth)**

Symbol Rates <sup>1, 10</sup>	Modulation and Encoding <sup>1, 10</sup>	Ranging	Coding LDPC <sup>4</sup>	Space Data Link Protocol AOS <sup>7</sup> , USLP <sup>5</sup>	Space Data link Security
128ksps ≤ symbol rate ≤ 4 Msps <sup>7</sup>	Filtered OQPSK + NRZ-L	No	Code Rates 1/2, 2/3, 4/5, 7/8, uncoded. Depending on the codeword selected, the following codeword size and ASM is to be used:	Depending on the codeword selected, the following AOS Frame size is used:	CCSDS Space Data Link Security Protocol <sup>8</sup>
64 ksps ≤ symbol rate ≤ 4 Msps <sup>7</sup>	Filtered BPSK + NRZ-L	No			
64 ksps < symbol rate ≤ 1.024Mpsps <sup>7</sup>	PCM/PM/Bi-phase-L (modulation on residual carrier)	Yes <sup>2</sup>	<ul style="list-style-type: none"> <li>• 4096 octets plus 64 bit ASM (for rate 1/2)</li> <li>• 3072 octets plus 64 bit ASM (for rate 2/3)</li> <li>• 2560 octets plus 64 bit ASM (for rate 4/5)</li> <li>• 1020 octets plus 32 bit ASM (for rate 7/8)</li> <li>• Uncoded size: 2048 octets plus a 32 bit ASM</li> </ul>	<ul style="list-style-type: none"> <li>• 2048 octets (for LDPC rates 1/2, 2/3, 4/5, or uncoded)</li> <li>• 892 octets (for LDPC rate 7/8)</li> </ul>	
0.1 ksps <sup>9</sup> ≤ symbol rate ≤ 64 ksps	PCM/PSK/PM + NRZ-L (modulation on subcarrier)	Yes <sup>2</sup>	LDPC Code rate 1/2 using the following codeword size and ASM: <ul style="list-style-type: none"> <li>• 256 octets plus 64 bit ASM</li> <li>• Uncoded size: 128 octets plus a 32 bit ASM</li> </ul>	<ul style="list-style-type: none"> <li>• 128 octets (for LDPC rate 1/2 or uncoded)</li> </ul>	



# Summary

- In this report, we have successfully verified the spectral and BER performance of the bi-phase filtering recommended by ESA:
  - Spectral performance with effective mod index  $m'$  of 1.0 and filter cut-off at  $f_{\text{cut}}=4.0R_s$
  - $E_b/N_0$  degradation at BER of  $1e-5$  for various combinations of  $m'$  and  $f_{\text{cut}}$  settings
- Our spectral results show that there is some small potential spectral violations for  $m'$  of 1.25 using  $f_{\text{cut}}$  of 3.5 $R_s$  therefore we reduced  $f_{\text{cut}}$  setting to 3.3 $R_s$ .
- For NASA missions, necessary BW for NTIA mask needs to be increased to cover the spectral emissions.
- Question: When PN ranging is on, does shaping the bi-phase pulse have any additional impact to chip tracking performance and also how does ranging impact BER of bi-phase detection?

