

SLS-RFM_23-02 MFSK for very low data rates

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- Background
- MFSK basics
- ESA study
- Way forward





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Background



Deep-space missions under certain conditions (high Doppler rate, high phase noise or very low SNR) challenging to demodulate using coherent schemes.

-> non-coherent *MFSK* (Multiple Frequency-Shift Keying) can be a solution for communication in these conditions.

NASA/JPL has already used MFSK:

- Mars rovers: EDL (DTE X-band)
- Juno: Deep-Space Manoeuvres and Jupiter Orbit Insertion
- Europa Clipper, ...?
- There is also a simpler "beacon tone" option to transmit basic spacecraft status

ESA has not yet used MFSK, but:

- Now implemented in next gen deep-space transponders
- Ongoing study:
 - analyse use cases and solutions
 - implement MFSK receiver
- DOR tones can be used as beacon tones
- **Other agencies?**

The topic has been for some time in the **RFM charter**:

18) Study modulation technique and position for pilot symbols of high order modulations

used in conjunction with the codes of CCSDS 131.0-B-2

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• Entry, Descent and Landing (high Doppler rate uncertainty)

Solar conjunction

(high phase and amplitude scintillation)

Safe/Survival mode

(very low SNR)

MFSK modulation



Classical MFSK: information resides in the tone frequency

Special MFSK: information resides in the frequency separation between carrier and tone



MFSK modulation



Classical MFSK: information resides in the tone frequency

Special MFSK: information resides in the frequency separation between carrier and tone



[MFSK signal from IDST transponder, OL recordings performed at ESOC, Dec 2022]

MFSK modulation

Modulation parameters:

- Modulation type [classical MFSK / special MFSK]
- Modulation order: M
- Tone duration: T_S
- Separation between tones: f_T
- Separation carrier to 1st tone: f_{T0}
- Tone waveform [sine / square] only applicable to special MFSK

In classical MFSK, separation between tones (f_T) must be sufficient to cover the maximum Doppler error In special MFSK, only the separation of the first tone (f_{T0}) must be higher than the maximum Doppler error



MFSK receiver architecture



- Signal recorded in open-loop, ideally Doppler-precompensated
 - Sampling frequency must cover all possible tones + Doppler uncertainty
 - Possibility to array multiple antennas to improve SNR -> potential cross-support
- Demodulation with quasi-real-time tool or offline tool
 - Depending on operational needs and processing power
 - Post-processing with finer Doppler correction might further improve detection
- Tone detection based on 2-D FFT
 - search for best frequency match and best frequency rate match





So far, JPL use of MFSK seems limited to indication of spacecraft status or events.

Can we use it to transmit TM according to CCSDS?

Coding:

- Codes in 131.0-B are designed for AWGN and coherent demodulation, we cannot expect same coding gain
- Lack of references of performance for 131.0-B codes with MFSK
- Frame duration for standard lengths become very long
 - Risk of losing a full frame might not be acceptable, in particular for EDL (7 minutes of terror!)

Data link and upper layers:

- Overhead increases transmission time
- But without frame structure, no standard way for cross-support (SLE)



- To our knowledge, no spacecraft has used MFSK for uplink so far
 - But there seems to be interest for future missions
- Applicable in the same scenarios as for downlink
- Simple to implement on ground station, but complexity is transferred to spacecraft
- Coding and upper layers?



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ESA MFSK study



- ESA-funded study, carried out by research institutes: CTTC + Ceit
- Inputs from previous ESA studies on EDL, solar conjunction
- Focused on downlink scenarios
- Main objectives:
 - Analysis of scenarios and selection of modulation <u>and coding</u> schemes
 - Implementation of MFSK receiver

ESA MFSK study - scenarios



Worst-case scenarios considered during study:

	EDL (X band)	Solar conjunction (X-band)	Solar conjunction <i>(Ka band)</i>	Safe/Survival mode (X band)
Doppler frequency error	<±50 kHz	<±100 Hz	<±400 Hz	<±100 Hz
Doppler frequency rate error	<±700 Hz/s	<±1.5 Hz/s	<6 Hz/s	<1.5 Hz/s
S/N0	> 14.2 dBHz	≥ 30 dBHz	≥ 30 dBHz	≥ 6 dBHz
Amplitude scintillation index		< 0.85 (Sun-Earth-Probe angle = 1 deg)	< 0.4 (Sun-Earth-Probe angle = 0.5 deg)	

Ka-band

during M2020 Entry, Descent, and

Landing,"]



Complexity

3.4e8 op/s

1.1e8 op/s

1.0e11 op/s

5.1e10 op/s

7.9e7 op/s

1.3e8 op/s

1.4e7 op/s

1.2e7 op/s

*

Optimisation done for the Modulation Symbol rate Tone Tone Scenario waveform Μ duration separation index (approx.) worst-case scenarios to achieve max symbol rate Classical 0.5 sps 128 1 kHz 14 s NA MFSK with acceptable Safe/survival mode Special computational MFSK 512 10 Hz 17 s 80 deg 0.5 sps complexity, while keeping sinewave P_{fd}<10⁻⁴ Classical 64 100 kHz 2.8 s NA 2 sps MFSK EDL ...this means correctly **Special** MFSK 512 200 Hz 2.9 s 70 deg 3 sps detecting **99.99%** of the sinewave tones. Classical 128 50 kHz 0.11 s NA 63 sps Maybe too harsh! MFSK Solar conjunction Special X-band MFSK 512 10 kHz 0.16 s 70 dea 56 sps For M2020 EDL. squarewave requirement was 90% Classical 128 10 kHz 0.04 s NA 160 sps MFSK [L. Mauger et al., "Direct to Earth Solar conjunction **Communications Using MFSK Tones Special**

MFSK

squarewave

512

1 kHz

0.06 s

70 deq

150 sps

Relaxation of the Pfd to 10² would allow increase to ~0.7sps (¹40% rate) with less complexity

Previous ESA study on EDL gave advantage to classical MFSK,but these results show opposite





[from ESA MFSK study]



For <u>coded signals</u>, soft demapper instead of hard output of FFT detector





[J. Gómez-Vilardebó, X. Mestre, M. Navarro, J. F. Sevillano, R. Abelló and J. Quintanilla, "Non-Coherent Receiver Design for MFSK Modulations in Deep Space Missions,"]





- Shortest LDPC frame with k=1024, $r=4/5 \rightarrow 1344$ symbols
 - 45 min at 0.5 sps (safe mode)
 - 7.5 min at 3 sps (EDL)
 - 21 s at 63 sps (solar conjunction X-band)
- Shorter frames possible with convolutional, but worse performance
- End-to-end simulations including coding are intensive for current Matlabbased simulator, preliminary results based on reduced scenario:

	Optimised scenario for safe mode	Reduced scenario
М	128	32
Ts	~14 s	3 s
N _{FFT}	~600000	9600
N _w	3	1
N _R	201	1



- Higher code rates are better!
 - \succ LDPC 4/5 would be selected for this case
- More representative results expected at end of study 1

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Way forward



- Wait for final results of ESA study... Fall 2023 or Spring 2024
- RFM WG position on adding MFSK to 401.0-B?
 - Only downlink or also uplink?
 - Should "beacon tones" also be included? (simpler version of MFSK)
- Coding and upper layers?
 - If no TM frames, should CCSDS define a new format for "very low rate information"?
 - Or investigate other codes outside 131.0-B?

References



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- X. Mestre, J. Gómez, M. Navarro, J. F. Sevillano, R. Abelló and J. Quintanilla, "Non-coherent Receivers for Low Data Rate Transmissions under Weak Solar Scintillation," 2022 9th International Workshop on Tracking, Telemetry and Command Systems for Space Applications (TTC), Noordwijk, Netherlands, 2022
- Ongoing ESA-funded study, contracted to CTTC and Ceit (Spain) "Multiple Frequency-Shift Keying Modem for Very Low Data Rates"